



Review

# A review of parameterizations for modelling dry deposition and scavenging of radionuclides

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## Abstract

This article aims at reviewing the state-of-the-science parameterizations for modelling dry deposition and scavenging of atmospheric tracers, with a focus on radionuclides. These parameterizations are key components of the numerical models that are used for environmental forecast. We present detailed models and parameterizations. Both are characterized by many uncertainties.

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*Keywords:* Dispersion models; Dry deposition; Wet scavenging; Parameterizations

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## 0. Introduction

Dry deposition and scavenging processes (rainout and washout) are key processes for the evolution of radionuclides in the atmosphere. The current state-of-the-science dispersion models (the so-called chemistry-transport models) used for radionuclides strongly rely on parameterizations for these processes.

There are however many sources of uncertainty. The first source is related to the meteorological fields, especially the rain intensity or the cloud characteristics (liquid water content and diagnosis) for scavenging. The second source is related to the microphysical description of the processes. The behaviour of radionuclides is strongly related to their chemical form as they may be released in the atmosphere as gases and particles. For instance, it is usually assumed that most of cesium is bound to particles (aerosols) while iodine is either bound to atmospheric aerosols or in gaseous form (elemental form and organic iodine  $\text{CH}_3\text{I}$ ). The partitioning between gaseous form and particles and the size distribution of aerosols strongly affect dry deposition and scavenging.

Wet scavenging is usually parameterized by  $dc/dt = -Ac$  with  $c$  the concentration and  $A$  the scavenging coefficient (in  $\text{s}^{-1}$ ). One usually distinguishes in-cloud scavenging (*rainout*) and below-cloud scavenging by raindrops (*washout*). It is often recognized that the wet scavenging of radionuclides is mainly related to particles (Chamberlain, 1991, p. 133) and is more important for rainout. The parameterizations are highly uncertain due to the difficulty of characterizing the aerosol distribution and the uncertainties for rain intensity.

Dry deposition is often applied as a boundary condition for vertical turbulent diffusion through  $K_z \nabla c \cdot \mathbf{n} = E - v_{\text{dep}} c$  with  $K_z$  the eddy coefficient,  $E$  the surface emission (possibly for resuspension) and  $v_{\text{dep}}$  the dry deposition velocity.  $\mathbf{n}$  is the unit vector oriented upwards. The deposition velocity depends on the surface meteorological fields, on the land use coverage and on the physical and chemical properties of the tracer.

Parameterization for  $A$  and  $v_{\text{dep}}$  are therefore required for 3D modelling. Many articles have already been devoted to these topics, ranging from theoretical studies focussed on detailed modelling to empirical parameterizations. The objective of this paper is to summarize the available parameterizations (to our knowledge), to underline the common characteristics and the differences and to investigate the uncertainties. The hope is that it may be useful for a modeller to have the synthesis of the majority of published data.

This article is structured as follows. We briefly summarize in Section 1 the key facts for the possible form of the radionuclides in the atmosphere. We emphasize on the measurements following the Chernobyl accident. Detailed and parameterized models are given in Sections 2 and 3 for dry deposition and wet scavenging, respectively. For each process, detailed models are presented and then empirical models, usually based on tuning to measurements. The focus is put on particles and wet scavenging. The article ends with conclusions.

## 1. Background for radionuclides

Three features have a strong impact for the loss properties of atmospheric radionuclides: the partitioning between gases and aerosols, the partitioning between the organic and inorganic forms and the aerosol size distribution.

*Gas/aerosol partitioning:* The radionuclides may be released into the atmosphere as gases and/or particles. The partitioning between both phases is a crucial issue for loss processes and there are many uncertainties reported in the literature, especially following the Chernobyl accident. For instance, short-range measurements (Ogorodnikov et al., 1994) indicate a gaseous fraction ranging from 30% to 90% of the total mass. Measurements in Great Britain (Clark and Smith, 1988) give a ratio of particle mass to gaseous mass ranging from  $\frac{1}{3}$  to  $\frac{1}{2}$ . This ratio is estimated to range from  $\frac{1}{5}$  to  $\frac{1}{3}$  in Germany (München, Chamberlain, 1991, p. 124), to

be of magnitude  $\frac{1}{4}$  in Sweden (Chamberlain, 1991, p. 84) and to vary with time. Baklanov and Sorensen (2001) estimate that 70% of the total release was in gaseous form during the first weeks (up to 90% of iodine emissions; see also Pöllänen et al., 1997). Other measurements indicate a gaseous fraction ranging from 50% to 85% in southern Finland (Jylhä, 1991).

Some theoretical studies have tried to compute the partitioning between both phases. For instance, in Budyka (2000), a condensation/evaporation model (unfortunately poorly detailed) gives the evolution of mass transfer between the gaseous and the particulate phases. The resulting gaseous fraction depends on the total atmospheric aerosol concentrations (10, 30, 100  $\mu\text{g m}^{-3}$ , standard values over Europe) and the size distribution (the diameters are in the range 0.015–0.4  $\mu\text{m}$ ). The resulting aerosol is centred around 0.5  $\mu\text{m}$  with a gaseous fraction of 15% (for a total atmospheric aerosol concentration of 100  $\mu\text{g m}^{-3}$ ), 35% (for 30  $\mu\text{g m}^{-3}$ ) and 60% (for 10  $\mu\text{g m}^{-3}$ ).

*Organic/inorganic partitioning:* Another source of uncertainties is related to the inorganic/organic partitioning for iodine. The dissolution properties are indeed not similar. For instance, the measurements in München (cited above) indicate a growth of the organic part (from 43% to 59% of the total mass, including gas and aerosols) due to the atmospheric ageing.

*Aerosol size distribution:* Besides the chemical composition, the loss processes strongly depend on the aerosol size distribution, through nonlinear processes (see the next sections). Hereafter,  $r_p$  will stand for the particle radius,  $d_p$  for the particle diameter.

For the Chernobyl accident, most authors reported that the aerosols were in the submicronic range ( $r_p < 1$  or 2  $\mu\text{m}$  in Brandt et al., 2002; Chamberlain, 1991, pp. 86 and 92). In Ogorodnikov et al. (1994), the size of the emitted particles differ according to the period.

In the first weeks, emissions are in the submicronic range (with an aerosol mean aerodynamic diameter (AMAD) or  $d_a$ , of 0.6–0.7  $\mu\text{m}$ ). We recall that for a particle density  $\rho_p \simeq 1.4 \text{ g cm}^{-3}$ ,  $d_a \simeq 1.2 \times d_p$ . This is in coherence with data from Baklanov and Sorensen (2001). Some typical values of particle distributions can be found for iodine and cesium in Table 1 (Baklanov and Sorensen, 2001). The aerosol distribution is a lognormal law of variance  $\sigma$ , centred around the AMAD. Jylhä (1991) reports a size distribution of [0.65–0.93]  $\mu\text{m}$  for cesium 137 and of [0.33–0.57]  $\mu\text{m}$  for particulate iodine. This reference also indicates that rain measurements in Switzerland

Table 1  
Particle distribution of cesium and iodine (Baklanov and Sorensen, 2001, following the Chernobyl accident)

Radionuclide	$v_{\text{dep}}$ ( $\text{cm s}^{-1}$ )	AMAD ( $\mu\text{m}$ )	$\sigma$
$^{137}\text{Cs}$	0.1	0.68	1.8–2.5
$^{134}\text{Cs}$	0.12	0.59	2–2.5
$^{131}\text{I}$	0.6	0.48	3–4
$^{133}\text{I}$	0.7	0.6	—

illustrate the bounds to “classical aerosols” (sulfates/nitrates). Sparmacher et al. (1993) report a measured aerosol distribution (in Germany) from 10 nm to 5  $\mu\text{m}$ , centred around 1  $\mu\text{m}$ .

In the second period (after the building of the sarcophagus), the AMAD was up to 6–7  $\mu\text{m}$  due to resuspension. Moreover, many studies (for instance Pöllänen et al., 1997) report coarse radioactive particles (“hot spots” beyond 20  $\mu\text{m}$ ) measured in Northern and Eastern Europe.

## 2. Dry deposition

### 2.1. Field measurement data

There are many references devoted to measurement data of dry deposition. For radionuclides, we focus here on particles. All the measured data indicate the sensitivity to the size distribution. Garland (2001) reviews some methods that are currently used (typically throughfall and eddy correlation methods). The measurements are usually characterized by a large uncertainty (Fig. 1). We refer to the discussion in Garland (2001) and to the review in Gallagher et al. (1999).

Garland (2001) reported a “mean” value  $v_{\text{dep}} = 0.5 \text{ cm s}^{-1}$  for cesium, derived from measurements following the Chernobyl accident. However, the value strongly depends on the natural surface: 0.05  $\text{cm s}^{-1}$  for grass, [0.07–0.55]  $\text{cm s}^{-1}$  for forest (the weakest value is for conifers). These values are coherent with those used in models with constant deposition velocities (see below). Notice that these constant values do not describe the strong sensitivity with respect to the size distribution (Fig. 1).

### 2.2. Detailed models

#### 2.2.1. For gases

Key references are Wesely (1989) and Wesely and Hicks (2000). The deposition process is usually

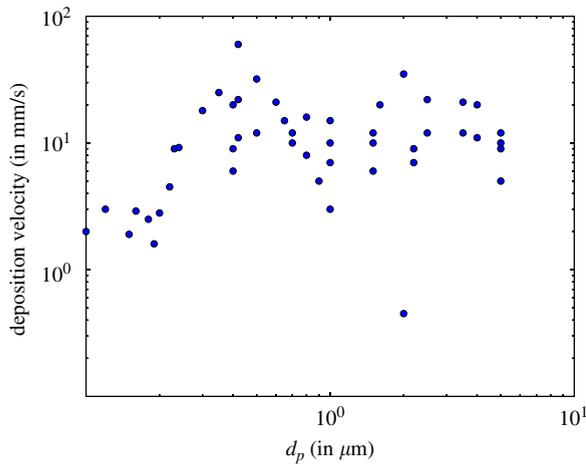


Fig. 1. Measured deposition of particles to forest (adapted from Gallagher et al., 1999).

interpreted in analogy to electrical resistance. The deposition to the surface is supposed to be controlled by three resistances in series: the aerodynamic resistance ( $R_a$ ), the quasi-laminar layer resistance ( $R_b$ ) and the surface resistance ( $R_s$ ). The total resistance (the inverse of the dry deposition velocity) is defined by the sum of these three resistances

$$v_{\text{dep}} = \frac{1}{R_a + R_b + R_s}. \quad (1)$$

The possible expression of the resistances can be found in classical textbooks (for instance Seinfeld and Pandis, 1998). For the sake of clarity, we do not enter into details.

### 2.2.2. For particles

For particles, the particle settling is supposed to operate in parallel with the previous processes. Moreover, one usually assumes that the surface resistance can be neglected because the particles adhere to the surface. With the electric resistance analogy, one gets (Seinfeld and Pandis, 1998):

$$v_{\text{dep}} = u_{\text{grav}} + \frac{1}{R_a + R_b + R_a R_b u_{\text{grav}}}, \quad (2)$$

where  $u_{\text{grav}}$  is the gravitational settling velocity (see below for the computation).

An alternative formula for the dry deposition velocity is (Venkatram and Pleim, 1999)

$$v_{\text{dep}} = \frac{u_{\text{grav}}}{1 - e^{-(R_a + R_b)u_{\text{grav}}}}. \quad (3)$$

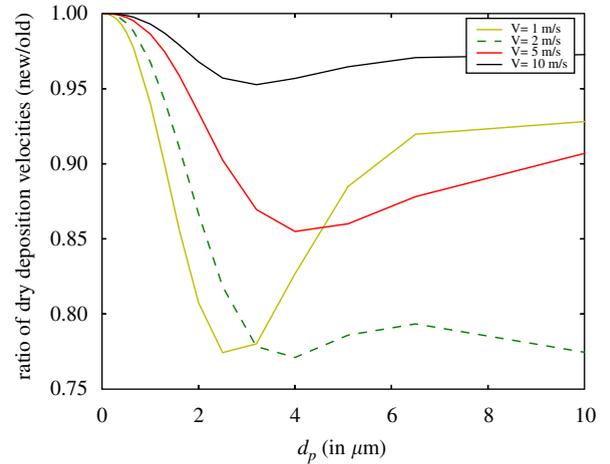


Fig. 2. Distribution of the ratio of dry deposition velocities (the new parameterization corresponds to Venkatram and Pleim (1999), the old parameterization to the classical formula) with respect to the aerosol diameter, for different values of wind velocity. The particle density is  $\rho_p = 1 \text{ g cm}^{-3}$ . The land use coverage is “needleleaf trees” in summer.

The advantage of this parameterization is that it satisfies mass conservation.

The ratio of the dry deposition velocity computed with Eq. (3) to this computed with Eq. (2) is given in Fig. 2 for aerosols of density  $1 \text{ g cm}^{-3}$ . The ratio is computed for different wind velocities as a function of the aerosol diameter. The computation of the resistances follows Zhang et al. (2001) and takes into account Brownian diffusion, inertial deposition and interception by ground.

Eq. (3) leads to smaller values of  $v_{\text{dep}}$  (the reduction is up to 20% for the largest particles for low wind velocities). Notice that the differences may be significant (as opposed to what is surprisingly reported in Venkatram and Pleim, 1999).

The parameterization of the dry deposition to water surface may also follow different laws. We refer for instance to Pryor et al. (1999) and Travnikov and Ilyin (2005).

*Gravitational settling velocity:* For small particles, say  $d_p \leq 20 \mu\text{m}$ ,  $u_{\text{grav}}$  is given by the Stokes formula (for instance Seinfeld and Pandis, 1998)

$$u_{\text{grav}} = \frac{d_p^2 (\rho_p - \rho_{\text{air}}) g C_c}{18 \mu_{\text{air}}}, \quad (4)$$

where  $\rho_p$  (in  $\text{kg m}^{-3}$ ) is the particle density,  $\mu_{\text{air}}$  the dynamic viscosity of air (in  $\text{Pa s}$ ) and  $g$  the gravity constant (in  $\text{m s}^{-2}$ ).  $C_c$  is the Cunningham correction factor that takes into account the slipping

effects for the finest particles:

$$C_c = 1 + \frac{2\lambda_{\text{air}}}{d_p} \left( 1.257 + 0.4 \exp\left(-0.55 \frac{d_p}{\lambda_{\text{air}}}\right) \right), \quad (5)$$

where  $\lambda_{\text{air}}$  is the air mean free path (in m):  $\lambda_{\text{air}} = 2\mu_{\text{air}}/(P\sqrt{8/\pi RT})$ , with  $R$  the gas constant (in  $\text{J K}^{-1} \text{kg}^{-1}$ ),  $T$  the temperature (in K) and  $P$  the pressure (in Pa).

For larger particles (or cloud droplets), the Stokes law is no longer valid and one needs to use a drag coefficient. A nonlinear system has to be solved (Seinfeld and Pandis, 1998):

$$u_{\text{grav}} = \sqrt{\frac{4gd_p C_c \rho_p}{3C_D \rho_{\text{air}}}}, \quad (6)$$

where the drag coefficient  $C_D$  is an analytical function of the Reynolds number of the particle (and, as such, a function of  $u_{\text{grav}}$ ). This equation can be solved by a Newton algorithm in few iterations. One also refers to Näslund and Thaning (1991). Notice that there exist many available parameterizations for the drag coefficient. The parameterization given in Näslund and Thaning (1991) has the advantage of providing a smooth function (as opposed to this of Seinfeld and Pandis, 1998).

### 2.3. Parameterized models

Some reference values for the dry deposition of iodine and cesium can be found in Table 2 (following Brandt et al., 2002). Additionally, the maximal values of dry deposition velocities are provided as a function of the land use coverage in Table 3.

It is sometimes advocated to relate the dry deposition velocities of radionuclides to the corresponding form of  $\text{SO}_x$ , that is to say to use  $\text{SO}_2$  as a typical gas-phase component and sulfate aerosols ( $\text{SO}_4^{2-}$ ) as a typical particle. For instance, for the EURAD model (Haas et al., 1990), the dry deposition velocity of cesium is half of this of aerosol sulfate. Baklanov and Sorensen (2001) reports the use of the dry deposition velocity of  $\text{SO}_2$  for gas-phase iodine ( $0.5 \text{ cm s}^{-1}$  for Baklanov and Sorensen, 2001) and the one of sulfate aerosols for particulate iodine ( $0.1 \text{ cm s}^{-1}$ ).

In Baklanov (1999, p. 11),  $0.05 \text{ cm s}^{-1}$  is the reference value for cesium 137. In Brandt et al. (2002), the default value for all species is  $0.2 \text{ cm s}^{-1}$ . In Raes et al. (1991), the reference is the dry deposition of accumulation mode of aerosols,  $0.1 \text{ cm s}^{-1}$ . For the modelling system RODOS

Table 2

Some values of dry deposition velocities (in  $\text{cm s}^{-1}$ ) from literature (from Brandt et al., 2002)

Reference	$^{131}\text{I}$	$^{137}\text{Cs}$
Hanna et al. (1991)	0.3	0.1
Maryon et al. (1992)	0.5	0.05
Klug et al. (1992)	0.15–2.0	0.1–0.5
Slinn and Slinne (1980)	—	0.31
Sehmel (1980)	0.1–2.0	0.04–0.5

Table 3

Values of dry deposition velocities (in  $\text{cm s}^{-1}$ ) (from Müller and Pröhl, 1993; Baklanov and Sorensen, 2001) according to the land use coverage (LUC) and for the RODOS system (Päsler-Sauer, 2003; Thykier-Nielsen et al., 1999)

LUC	Particulate	Elemental iodine	Organic iodine
Soil	0.05	0.3	0.005
Grass	1.05	0.15	0.015
Trees	0.5	5	0.05
Other plants	0.2	2	0.02
RODOS	0.1	0.8	0.01
RIMPUFF	0.1	1	0.05

(Päsler-Sauer, 2003, p. 16), some typical values are given for dry deposition over lawn with a wind at 10 m above the ground of  $4 \text{ m s}^{-1}$ :  $0.8 \text{ cm s}^{-1}$  for elemental iodine,  $0.01 \text{ cm s}^{-1}$  for organic iodine and  $0.1 \text{ cm s}^{-1}$  for aerosols. The RIMPUFF model (a component of RODOS, Thykier-Nielsen et al., 1999, pp. 20–23) indicates values of  $1 \text{ cm s}^{-1}$  for elemental iodine,  $0.05 \text{ cm s}^{-1}$  for organic iodine and  $0.1$  for aerosols. These values are summarized in Table 3.

We also refer to Garland (2001) for a discussion of the gap between model results and observational data. A possible explanation is that the models may omit electrical effects and thermo- and diffusion-phoretic forces.

In Quélo et al. (2006), we have used a robust approach by taking  $0.2 \text{ cm s}^{-1}$  for iodine and  $0.2$  or  $0.1 \text{ cm s}^{-1}$  for cesium (with model-to-data comparisons for the Chernobyl and Algeciras cases).

A common feature of these parameterizations is that the dry deposition velocity of elemental iodine is greater than the one of aerosols, which is greater than the one of organic iodine.

### 3. Wet scavenging

Wet scavenging is usually partitioned in *in-cloud* scavenging and *below-cloud* scavenging. Some

authors advocate rather to distinguish *nucleation scavenging* and *impaction scavenging* in order to give a more accurate process description (see the discussion in Tost et al., 2006). Even if rainout is a key process, we focus on below-cloud scavenging because there are not a lot of data related to in-cloud scavenging.

### 3.1. Field measurement data

The experimental data indicate a strong sensitivity of the scavenging coefficient with respect to the particle size. The scavenging is much stronger for particles below, let say, 0.3 μm or above 1–2 μm than for particles in the so-called Greenfield gap (typically in the range [0.3–1] μm, Greenfield, 1957). The measured data are highly “scattered” (Fig. 3 as an example) and the model-to-data comparison is usually difficult to perform.

Many data are the basis of the fitting presented in Section 3.3 (for instance Sparmacher et al., 1993).

### 3.2. Detailed models

A crucial issue is to fix the distribution of the raindrops and the resulting falling velocities. We first describe models of rain and then the scavenging of gases and particles.

#### 3.2.1. Representation of the rain

In the following,  $p_0$  is the rain intensity (in mm h<sup>-1</sup>).  $D_r$  (in meter) stands either for the

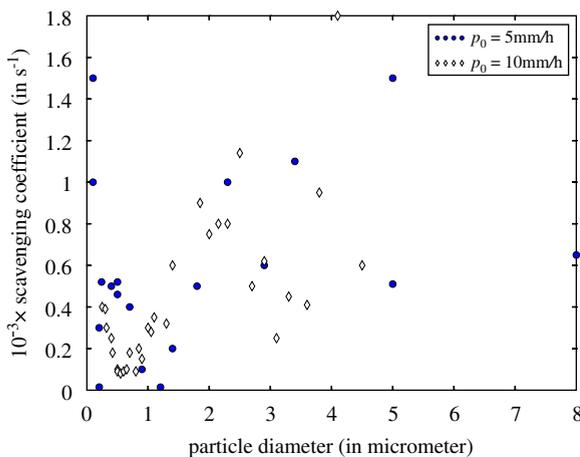


Fig. 3. Measured scavenging coefficient for two rain intensities ( $p_0 = 5$  and  $10 \text{ mm h}^{-1}$ ), from Baklanov and Sorensen (2001) and the references therein. These values are given as illustration for the “scattering”.

raindrop diameter of a monodisperse distribution or for the representative diameter of a polydisperse distribution. Other units are always mentioned.

3.2.1.1. *Raindrop distribution.* The raindrop distribution is usually described by a Gamma function with four parameters

$$n_r(D_r) = \alpha_0 D_r^\alpha \exp(-\beta D_r^\gamma). \tag{7}$$

The two classical cases are the Marshall–Palmer distribution ( $\alpha = 0$ ;  $\gamma = 1$ ) and the Khrgian–Mazin distribution ( $\alpha = 2$ ,  $\gamma = 1$ ). We refer to Table 4. It is usually recognized that the Marshall–Palmer distribution leads to an overestimation of the finest droplets.

de Wolf (2001) proposes a distribution based on a fit to measurements as a function of the rain intensity

$$n_r(D_r) = \alpha(p_0) \times 2.6 \times 10^{-6} p_0^{-0.384} D_r^{2.93} \times \exp(-2.69 p_0^{-0.186} D_r), \tag{8}$$

with the normalization factor  $\alpha(p_0) = 1.047 - 0.0436 \ln p_0 + 7.34 \times 10^{-3} (\ln p_0)^2$ . We refer to the discussion in Henzing et al. (2006) for a comparison of the different parameterizations and the impact for below-cloud scavenging of particles.

In Mircea et al. (2000), lognormal distributions are proposed on the basis of measurements.

3.2.1.2. *Representative diameter.* There exist several parameterizations for the representative diameter of raindrops of the form  $D_r = A p_0^B$  (see Table 5, in mm there):

$$D_r = [0.2431 - 0.97] p_0^{[0.158-0.25]}. \tag{9}$$

The comparison between the parameterizations is given in Fig. 4. One notices the spread of results, especially for large rain intensities. As already said, the Marshall–Palmer distribution leads to an over-

Table 4

Typical ranges for raindrop diameters given by the Marshall–Palmer and Khrgian–Mazin distributions (Mircea and Stefan, 1998)

Rain type	$p_0$ (mm h <sup>-1</sup> )	Range for $D_r$ (mm)
Weak	[1–5]	[0/0.001–0.1]
Moderate	[5–100]	[0/0.01–1]
Strong	[100–500]	[0/0.1–10]

In the notation  $[a/b - ]$ ,  $a$  stands for the lower value of the Marshall–Palmer distribution,  $b$  for the one of the Khrgian–Mazin distribution.

estimation of finest droplets (which results in the overestimation of collision efficiencies and then of scavenging; see below).

Table 5  
Some parameterizations for the representative diameter of raindrops,  $D_r$  (here in mm)

Index	Reference	$D_r$ (mm)
1	Pruppacher and Klett (1998, p. 34)	$0.976 \times p_0^{0.21}$
2	Marshall–Palmer	$0.243 p_0^{0.21}$
3	Andronache (2004)	$0.24364 p_0^{0.214}$
4	Loosmore and Cederwall (2004)	$0.97 p_0^{0.158}$
5	Mircea et al. (2000)	$[0.63–0.72] p_0^{0.23}$
6	Underwood (2001, p. 35)	$0.7 p_0^{0.25}$

Mircea et al. (2000) uses measurements over the Eastern Mediterranean area (which results in a range and not in a given value).

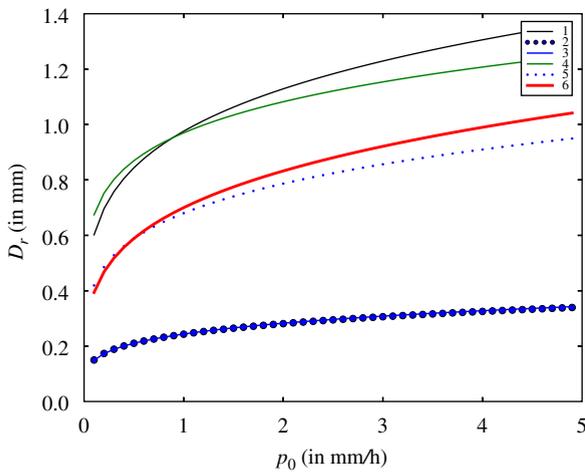


Fig. 4. Evolution of the representative diameter as a function of the rain intensity for different parameterizations. The index is given in Table 5.

Table 6  
Some parameterizations for the falling velocity of raindrops (in  $\text{m s}^{-1}$ )

Index	Reference	$U_{\text{drop}}$ ( $\text{m s}^{-1}$ )
1	Kessler (Andronache, 2003, p. 143; Mircea and Stefan, 1998, Table 2)	$130\sqrt{D_r}$
2	Seinfeld (1985, p. 632)	$9.58 \left[ 1 - \exp\left(-\left(\frac{D_r}{0.171 \times 10^{-2}}\right)^{1.147}\right) \right]$
3	Seinfeld and Pandis (1998) and Mircea et al. (2000)	Terminal velocity
4	Andronache (2004)	$3.778 \times 10^3 D_r^{0.67}$
5	Loosmore and Cederwall (2004)	$4.854 D_r \exp(-195 \times 10^{-3} D_r)$

$D_r$  is in meter.

3.2.1.3. *Falling velocity.* The falling velocity  $U_{\text{drop}}$  (in  $\text{m s}^{-1}$ ) can be computed as a function of diameters through different parameterizations (Table 6). For the terminal velocity, due to the size of falling raindrops (of diameter bigger than  $20 \mu\text{m}$ ), the Stokes formula (4) cannot be used and a system similar to (6) has to be used

$$U_{\text{drop}} = \sqrt{\frac{4gD_r C_c \rho_{\text{water}}}{3C_D \rho_{\text{air}}}} \quad (10)$$

Comparisons among the parameterizations are given in Fig. 5 with the parameterization “4” of the representative diameter (index = 4 of Table 5). The discontinuity for the index 3 is related to the change of the computation of the drag coefficient for particle Reynolds number greater than 500.

### 3.2.2. Scavenging of gases

Notice that these detailed models are already parameterized versions of microphysical models (see for instance Sportisse and Du Bois, 2002; Sportisse and Djouad, 2003 for the justifications).

3.2.2.1. *Below-cloud scavenging (washout).* We refer for instance to Sportisse and Du Bois (2002). The scavenging coefficient for a gas is given by

$$A = \frac{6LD_g Sh}{D_r^2} \exp\left(\frac{6D_g Sh z}{D_r^2 U_{\text{drop}} HRT}\right), \quad (11)$$

with  $L$  the liquid water content for rain (volume of falling water per volume of air),  $D_g$  the molecular gas-phase diffusivity for the scavenged gas,  $H$  the Henry’s coefficient (the efficient coefficient if the acidity of the rain is taken into account).  $Sh$  is the Sherwood number:

$$Sh = 2 + 0.6 \left(\frac{U_{\text{drop}} D_r}{v_{\text{air}}}\right)^{1/2} \left(\frac{v_{\text{air}}}{D_g}\right)^{1/3}, \quad (12)$$

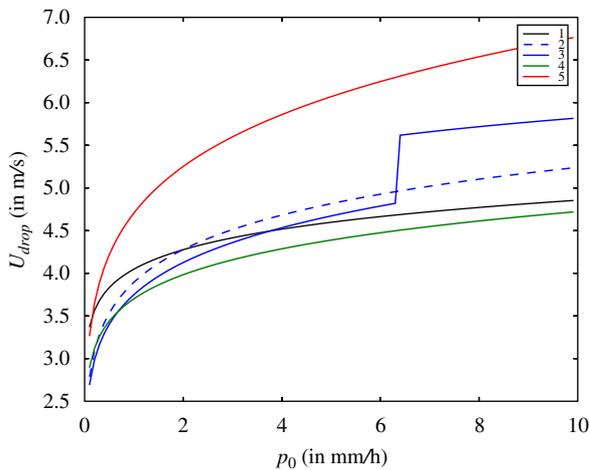


Fig. 5. Evolution of the falling velocity of raindrops with respect to rain intensity. The index is related to the parameterization of falling velocity. The representative diameter is computed with the parameterization “4” for  $D_r$ .

with  $\nu_{\text{air}} = \mu_{\text{air}}/\rho_{\text{air}}$  the kinematic viscosity of air (in  $\text{m}^2 \text{s}^{-1}$ ).  $z$  is the distance of the falling raindrop from the cloud.

**3.2.2.2. In-cloud scavenging (rainout).** In-cloud scavenging is much more difficult to parameterize especially because aqueous-phase chemistry has to be taken into account. We refer for instance to Sportisse and Djouad (2003) for a rigorous derivation of kinetic mass transfer coefficients. We also refer to Tost et al. (2006) for the description of a detailed coupled model between the gas phase and the aqueous phase.

A simplified model is given in Roselle and Binkowski (1999)

$$A = -\frac{1 - e^{-\tau_{\text{cld}}/\tau_{\text{washout}}}}{\tau_{\text{cld}}}, \quad (13)$$

where  $\tau_{\text{cld}}$  (in second) is the 3D timestep of the dispersion model if the cloud size exceeds the cell dimension, and is equal to 1 h otherwise.  $\tau_{\text{washout}}$  is the time required for the volume of water to precipitate to the ground. For a cloud of liquid water content  $L$  (in volume per volume), of depth  $\Delta z_{\text{cld}}$  and of total (air) volume  $V$ ,  $\tau_{\text{washout}}$  depends on the precipitation rate  $p_0$  (in  $\text{m s}^{-1}$ ) through  $LV = \tau_{\text{washout}} p_0 S$ , where  $S$  is the ground surface covered by the cloud. If we assume that  $V = S \Delta z_{\text{cld}}$ , one gets  $\tau_{\text{washout}} = L \Delta z_{\text{cld}} / p_0$ .

### 3.2.3. Scavenging of particles

The aerosol radius and diameter for a monodisperse distribution are  $r_p$  and  $d_p$  (in  $\mu\text{m}$ ),

respectively. In the following, we focus on the washout of particles (*below-cloud scavenging*), which corresponds to the scavenging for falling raindrops.

**3.2.3.1. Scavenging coefficient for a monodisperse distribution.** The scavenging rate of particles of diameter  $d_p$  is given by

$$\frac{dn_p(d_p)}{dt} = -A(d_p)n_p(d_p). \quad (14)$$

The scavenging coefficient  $A(d_p)$  is computed on the basis of the number of particles collected per time unit by raindrops. For a monodisperse distribution of raindrops, of representative diameter  $D_r$ , this gives

$$A(d_p) = \frac{\pi}{4} D_r^2 U_{\text{drop}}(D_r) E(D_r, d_p) N_r, \quad (15)$$

with  $E(D_r, d_p)$  the collision efficiency and  $N_r$  the total density of raindrops (in  $\text{m}^{-3}$ ). The collision efficiency is defined as the fraction of the particles of diameter  $d_p$  in the collision volume of the raindrop of diameter  $D_r$  that are actually scavenged. This formula assumes that the raindrop falling velocity is much bigger than the settling velocity of aerosols and that the representative diameter for raindrops is also much bigger than the aerosol diameter, which is met in practice.

As the rain intensity  $p_0$  (here in  $\text{m s}^{-1}$ ) is defined by

$$p_0 = \int_0^\infty \frac{\pi}{6} D_r^3 U_{\text{drop}}(D_r) n_r(D_r) dD_r \quad (16)$$

one gets in the monodisperse case  $p_0 = (\pi/6) D_r^3 U_{\text{drop}}(D_r) N_r$ . The scavenging coefficient is therefore

$$A(d_p) = \frac{3 E(D_r, d_p) p_0}{2 D_r}. \quad (17)$$

**3.2.3.2. Scavenging coefficient for a polydisperse distribution.** Let us consider a polydisperse distribution for raindrops, given by  $n_r(D_r)$  (in  $\text{m}^{-3} \text{m}^{-1}$ ). The scavenging coefficient is then integrated over the raindrop distribution

$$A(d_p) = \int_0^\infty \frac{\pi}{4} D_r^2 U_{\text{drop}}(D_r) E(D_r, d_p) n_r(D_r) dD_r. \quad (18)$$

A special attention has to be paid to the numerical integration due to the vanishing part of the distribution (tests not reported here).

3.2.3.3. *Collision efficiency.* A crucial point is to parameterize the collision efficiency  $E$ . The collision efficiency is governed by Brownian diffusion (in favour of small particles), inertial collision (for heavy particles) and interception (for large particles). The Brownian diffusion justifies that scavenging may be large for small particles (with diameters less than  $0.01\ \mu\text{m}$ ) while the inertial effect justifies that scavenging may be large for coarse particles (with diameters bigger than  $2\ \mu\text{m}$ ). The aerosols in the *Greenfield Gap* or *scavenging gap* (with diameters typically in the range  $[0.01; 2]\ \mu\text{m}$ ) are more weakly scavenged. Notice that this gap is not always observed due to neglected effects (see below).

The application of the Buckingham  $\pi$  theorem gives  $E$  as a function of five dimensionless parameters: the raindrop Reynolds number ( $Re = D_r U_{\text{drop}}/2\nu_{\text{air}}$ ), the particle Schmidt number ( $Sc = \nu_{\text{air}}/D_B$  with  $D_B = kTC_c/3\pi\mu_{\text{air}}d_p$  the particle Brownian diffusivity coefficient in  $\text{m}^2\ \text{s}^{-1}$  and  $k$  the Boltzmann constant in  $\text{JK}^{-1}$ ), the particle Stokes number ( $St = 2u_{\text{grav}}/g \times (U_{\text{drop}} - u_{\text{grav}})/D_r$ ) and the ratio of diameters ( $\phi = d_p/D_r$ ) and viscosities ( $\omega = \mu_{\text{water}}/\mu_{\text{air}}$ ).

$E$  is then computed by

$$E = \underbrace{\frac{4}{Re Sc} (1 + 0.4 Re^{1/2} Sc^{1/3} + 0.16 Re^{1/2} Sc^{1/2})}_{\text{Brownian diffusion}} + \underbrace{4\phi(\omega^{-1} + [1 + 2 Re^{1/2}]\phi)}_{\text{interception}} + \underbrace{((St - S^*)/(St - S^* + 2/3))^{3/2} (\rho_p/\rho_{\text{water}})^{1/2}}_{\text{impaction}}, \tag{19}$$

with the critical Schmidt number  $S^* = \frac{1.2+1/12\ln(1+Re)}{1+\ln(1+Re)}$ . The three terms correspond to Brownian diffusion, interception and impaction, respectively.  $E(D_r, d_p)$  and the different contributions are illustrated in Fig. 6 for a raindrop diameter  $D_r = 0.1\ \text{mm}$ .

The collision efficiency is a function of both the raindrop diameter and the aerosol diameter (Fig. 7). The spread in the results with respect to the raindrop diameter has to be compared with the uncertainties in the representative diameter (Fig. 4).

3.2.3.4. *Impact of a polydisperse distribution of aerosols.* Mircea et al. (2000) take into account the polydisperse nature of aerosols and raindrops. The aerosol distribution is given by a lognormal law. This results in an affine parameterization of  $A$

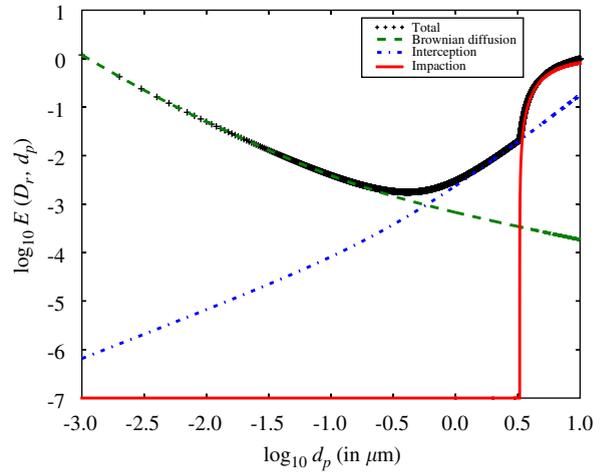


Fig. 6. Contributions of Brownian diffusion, interception and impaction for  $E(D_r, d_p)$ . The representative diameter for raindrops is  $D_r = 0.1\ \text{mm}$ .

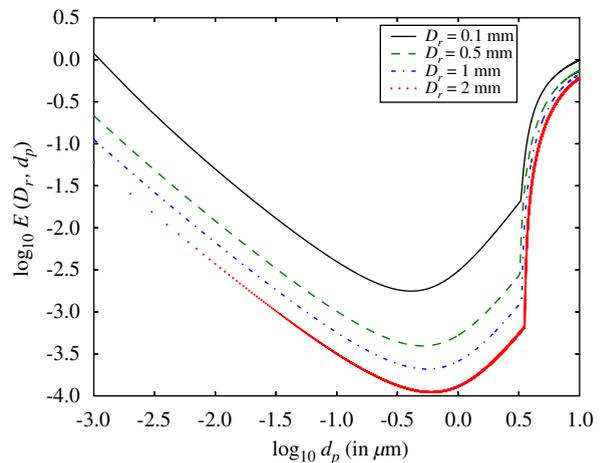


Fig. 7. Distribution of  $E(D_r, d_p)$  with respect to  $d_p$  for representative diameters  $D_r = 0.1, 0.5, 1$  and  $2\ \text{mm}$ .

as a function of rain intensity,  $A = a + bp_0$ . ( $a, b$ ) are given as function of the aerosol type (rural, urban, marine):  $a \in [1.58 \times 10^{-2} - 1.98]$  and  $b \in [2.17 \times 10^{-3} - 3.19 \times 10^{-1}]$ . The scavenging in an urban environment is much greater (one order of magnitude) than the scavenging in a remote environment (the maximal values correspond to the urban aerosol, the minimal values to the maritime aerosol). The main conclusion is also that the scavenging of a polydisperse distribution of aerosols is much greater than the scavenging of a monodisperse distribution (one order of magnitude at least). Moreover, there is a weak sensitivity with respect to the raindrop distribution.

We also refer to [Zhao and Zheng \(2006\)](#) for the investigation of appropriate numerical methods applied to the scavenging of polydisperse distributions (both for raindrops and aerosols, with lognormal distributions in the case study). Such methods are based on Monte Carlo techniques.

### 3.3. Parameterized models

#### 3.3.1. Justification of the parameterization $\Lambda = Ap_0^B$

For a monodisperse aerosol distribution with a representative raindrop diameter  $D_r$ ,  $\Lambda = 1.5Ep_0/D_r$ . As  $p_0$  is usually in  $\text{mm h}^{-1}$ , one has to apply a conversion factor  $10^{-3}/3600$ . With the spread in the estimation of  $D_r$ , Eq. (9), one gets

$$\Lambda \simeq [0.43 - 1.71] \times 10^{-3} Ep_0^{[0.75-0.842]}. \quad (20)$$

For  $E \in [0.1 - 1]$ , this leads to

$$\Lambda \simeq [4.3 \times 10^{-5} - 1.71 \times 10^{-3}] p_0^{[0.75-0.842]}. \quad (21)$$

In [Mircea and Stefan \(1998\)](#), it is proven that the scavenging coefficient (integrated over the raindrop distribution) can be put under the form  $\Lambda = Ap_0^B$ . The assumptions are a  $\Gamma$  function for the raindrop distribution and a constant collision efficiency.  $B$  is a function of the parameters of the  $\Gamma$  distribution and of the raindrop falling velocity (supposed to be under the form  $U_{\text{drop}}(D_r) \propto D_r^z$ ).

From [Table 4](#), this leads to

$$\Lambda \simeq \begin{cases} [0.753 - 0.875] Ep_0^{[0.78-0.86]} & \text{for strong rains,} \\ [1.91 - 1.95] Ep_0^{[0.78-0.86]} & \text{for moderate rains,} \\ [20.56 - 26.67] Ep_0^{[0.78-0.86]} & \text{for weak rains.} \end{cases} \quad (22)$$

Here,  $\Lambda$  is in  $\text{h}^{-1}$  and  $E$  (collision efficiency) has a constant value. With  $E \simeq 0.1$  and after a conversion to  $\text{s}^{-1}$ , one gets  $\Lambda \simeq 2.3 \times 10^{-5} p_0^{0.8}$  for heavy rains and  $\Lambda \simeq 6.6 \times 10^{-4} p_0^{0.8}$  for weak rains.

A similar approach is used in [Andronache \(2003\)](#) in order to justify the parameterization  $\Lambda = Ap_0^B$  with a Marshall–Palmer distribution of raindrops intensity and a polydisperse aerosol distribution ([Tables 7 and 8](#)).

#### 3.3.2. Empirical models derived from measurements

**3.3.2.1. Measurements after the Chernobyl accident.** [Jylhä \(1991\)](#) reports radar measurements in southern Finland after the Chernobyl accident. For cesium and iodine, this leads to  $\Lambda \simeq [10^{-5} - 10^{-3}] p_0^{[0.5-0.7]}$  with an average value  $10^{-4} p_0^{0.64}$

([Table 9](#)). Notice that the rain intensity was weak (less than  $1 \text{ mm h}^{-1}$ ).

**3.3.2.2. Measurements of ultrafine particles.** [Laakso et al. \(2003\)](#) report results of measurements for 6 years over forests in southern Finland. Five-hundred and eighty-eight hours of rain have been measured with rain intensities ranging up to  $20 \text{ mm h}^{-1}$  with an average value of  $0.9 \text{ mm h}^{-1}$ .

The scavenging coefficient has a minimal value for a diameter of  $0.2 \mu\text{m}$ . Scavenging coefficients range from  $7 \times 10^{-6}$  to  $4 \times 10^{-5}$  and have greater values than theoretically expected. A parameterized model is proposed for ultrafine particles ( $[10-510] \text{ nm}$ ) for rain intensities in the range  $[0-20] \text{ mm h}^{-1}$

$$\log \Lambda(d_p, p_0) = a + \frac{b}{\log d_p^4} + \frac{c}{\log d_p^3} + \frac{d}{\log d_p^2} + \frac{e}{\log d_p} + 0.24\sqrt{p_0}. \quad (23)$$

Up to first-order,  $\Lambda \sim \exp(0.24\sqrt{p_0})$ .

One also refers to [Andronache et al. \(2006\)](#) for a model evaluation.

#### 3.3.3. Parameterizations $\Lambda = Ap_0^B$

We have already justified the parameterization  $\Lambda = Ap_0^B$ . A typical value of  $B$  is 0.8. We have reported in [Table 10](#) of some values collected in the literature.

Table 7

Fitting of polydisperse raindrop and aerosol distributions ([Andronache, 2003](#))

$A$	$B$	Reference	Aerosol type
$6.67 \times 10^{-5}$	0.7	<a href="#">Andronache (2003)</a>	Urban
$1.28 \times 10^{-4}$	0.7	id.	Remote continental
$1.39 \times 10^{-4}$	0.7	id.	Marine
$1.28 \times 10^{-4}$	0.7	id.	Rural
$1.89 \times 10^{-4}$	0.7	id.	Free troposphere
$9.44 \times 10^{-5}$	0.7	id.	Polar
$2.44 \times 10^{-4}$	0.7	id.	Desert
$2.22 \times 10^{-4}$	0.7	id.	Marine
$8.33 \times 10^{-5}$	0.7	id.	Marine
$1.94 \times 10^{-4}$	0.7	id.	Dust
$1.00 \times 10^{-4}$	$[0.67-0.76]$	id.	Exp. data
$3.50 \times 10^{-4}$	0.78	id.	In-cloud

Coarse fraction of the aerosol distribution (diameter above  $10 \mu\text{m}$ ).

Table 8  
Fitting of polydisperse raindrop and aerosol distributions (Andronache, 2003)

A	B	Reference	Aerosol type
$2.34 \times 10^{-7}$	0.59	Sparmacher et al. (1993)	Exp. ( $d_p = 0.23$ )
$3.14 \times 10^{-7}$	0.60	id.	Exp. ( $d_p = 0.46$ )
$2.56 \times 10^{-7}$	0.94	id.	Exp. ( $d_p = 0.98$ )
$1.72 \times 10^{-6}$	0.61	id.	Exp. ( $d_p = 2.16$ )
$6.90 \times 10^{-6}$	0.92	Julya (1999)	Exp. (radionuclides)
$[2.36 \times 10^{-7} - 1.4 \times 10^{-6}]$	$[0.59-0.61]$	Andronache (2003)	Marine
$[2.78 \times 10^{-8} - 3.89 \times 10^{-8}]$	0.59	id.	Marine
$2.36 \times 10^{-7}$	0.59	id.	Mountains

Submicronic fraction of the aerosol distribution. “Exp.” stands for “experimental data”.

Table 9  
Parameterizations by fitting to observational data (Jylhä, 1991) after the Chernobyl accident

Species	A	B
$^{131}\text{I}(p)$	$7 \pm 5 \times 10^{-5}$	$0.69 \pm 0.12$
$^{133}\text{I}(p)$	$1.6 \pm 3 \times 10^{-5}$	$0.5 \pm 0.2$
$^{134}\text{Cs}(p)$	$2.8 \pm 0.6 \times 10^{-5}$	$0.51 \pm 0.07$
$^{137}\text{Cs}(p)$	$3.4 \pm 0.9 \times 10^{-5}$	$0.59 \pm 0.08$

Belot et al. (1988) advocate  $A = 4 \times 10^{-5}$  and  $B = 0.6$  for gaseous iodine. These values have been validated by other measurements (Caput et al., 1993) with an average value  $A \simeq 8.2 \times 10^{-5}$ .

For particulate iodine, the values are generally given on the basis of experimental data due to the underestimation provided by the theoretical values (related to the electrical interactions between liquid droplets and aerosols that are usually neglected).

CEC (1994) advocates the following values:

- for reactive iodine (probably “elemental”):  $A \simeq 10^{-4}$ ;
- for organic iodine:  $A \simeq 10^{-6}$ ;
- for submicronic particles ( $r_p < 1$ ):  $A \simeq 5 \times 10^{-5}$ ;
- for particles  $r_p = 5$ :  $A \simeq 10^{-4}$ ;
- for coarse particles  $r_p = 10$ :  $A \simeq 5 \times 10^{-4}$ .

The parameterization of the NAME model (Maryon et al., 1992) takes into account the different types of scavenging (Table 11). For rainout, dynamic rain and convective rains are splitted (which is usually the case in the outputs of meteorological models). The increase in the scavenging coefficient for convective rain is also taken into

account by other parameterizations (see Loosmore and Cederwall, 2004 and below).

A model based on measurement with a more complicated form is proposed in Baklanov and Sorensen (2001), with a focus on fine particles:

$$A(p_0, r_p) = \begin{cases} 8.4 \times 10^{-5} p_0^{0.79} & \text{if } r_p < 1.4 \mu\text{m}, \\ g(r_p)f(p_0) & \text{if } 1.4 < r_p < 10 \mu\text{m}, \\ f(p_0) & \text{if } r_p > 10 \mu\text{m}, \end{cases} \quad (24)$$

with  $f(p_0) \simeq 2.7 \times 10^{-4} p_0 - 3.62 \times 10^{-6} p_0^2$  and  $g(r_p) \simeq -0.15 + 0.32 r_p - 3.0 \times 10^{-2} r_p^2 + 9.34 \times 10^{-4} r_p^3$ . A specific point is that it takes into account scavenging for the smallest particles (as opposed to Näslund and Holmström, 1993). The results obtained with this parameterization are given in Fig. 8. Notice that the general behaviour is different from the classical one and that there is no Greenfield gap.

For in-cloud scavenging, ENVIRON (2005) uses  $A = 4.2 \times 10^{-4} p_0^{0.79}$  and is derived on basis similar to those used for the particles (see above) with a collision efficiency  $E \simeq 0.9$  and a raindrop diameter given by a law similar to the first case in Table 5 ( $D_r = 0.9 \times 10^{-4} p_0^{0.21}$  in mm Scott, 1978).

### 3.3.4. Other parameterizations

3.3.4.1. Parameterization based on the relative humidity. Brandt et al. (2002), following Pudykiewicz (1989), advocate a parameterization based on the relative humidity (RH). The advantages are that the rain intensity is usually highly uncertain and that it may describe in-cloud scavenging (which has the greatest impact for submicronic particles). When

Table 10

Some values of  $(A, B)$  for the parameterizations  $A = Ap_0^B$ 

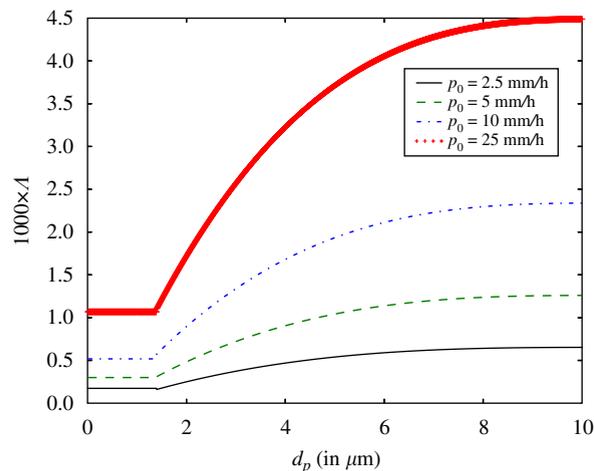
Reference	Default	Particulate	Elem. iodine	Org. iodine
Raes et al. (1991)	$(5 \times 10^{-5}; 0.8)$	—	—	—
Päsler-Sauer (2003)	—	$(8 \times 10^{-5}; 0.8)$	$(8 \times 10^{-5}; 0.6)$	$(8 \times 10^{-7}; 0.6)$
Belot et al. (1988) and Caput et al. (1993)	$8.2 \times 10^{-5}$ (gas)	$(4 \times 10^{-5}; 1)$ if $r_p < 1$ $(1.3 \times 10^{-4}; 1)$ if $r_p > 1$	$(4 \times 10^{-5}; 0.6)$	—
CEC (1994)	—	$5 \times 10^{-5}$ if $r_p < 1$ $10^{-4}$ if $r_p = 5$ $5 \times 10^{-4}$ if $r_p = 10$	$10^{-4}$	$10^{-6}$

For constant parameterizations, there is only one value.  $r_p$  is in  $\mu\text{m}$ .

Table 11

Parameterizations of the NAME model for rainout/washout (Maryon et al., 1992)

Scavenging	$A$	$B$
Washout	$8.4 \times 10^{-5}$	0.79
Convective rainout	$3.35 \times 10^{-4}$	0.79
Dynamic rainout	$8.4 \times 10^{-5}$	0.79

Fig. 8. Parameterizations of Baklanov and Sorensen (2001).  $A$  is multiplied by 1000.RH is above a threshold  $\text{RH}_t$ ,

$$A = A \frac{\text{RH} - \text{RH}_t}{\text{RH}_s - \text{RH}_t}, \quad (25)$$

with  $A = 3.5 \times 10^{-5}$ . In Brandt et al. (2002),  $\text{RH}_t = 80\%$  and  $\text{RH}_s = 100\%$ .

3.3.4.2. *Parameterization of electrical effects.* Taking into account the electrical effects may increase the scavenging coefficient up to a factor of 10 (Tripathi and Harrison, 2001). An additional term for the collision efficiency is proposed in Andronache (2004):

$$E_{\text{es}} = \frac{16KC_c Q_r q_p}{3\pi\mu_{\text{air}} U_{\text{drop}} D_r^2 d_p}, \quad (26)$$

with  $K = 9 \times 10^9$  (in  $\text{Nm}^2\text{C}^{-2}$ ),  $Q_r$  and  $q_p$  are the charges (in Coulomb, C) of a raindrop and of a particle, respectively. A parameterization with respect to the size is proposed

$$Q_r = \alpha x D_r^2, \quad q_p = \alpha x d_p^2, \quad (27)$$

with  $a = 0.83 \times 10^{-6}$  and  $\alpha$  a parameter that depends on the electrical environment (from the microphysical and meteorological states). The minimal value of  $\alpha$  is 2, the average value is 7.

3.3.4.3. *Parameterization of violent rains.* Loosmore and Cederwall (2004) advocate a correction to the existing parameterizations in order to describe the increase of the particle scavenging during convective rains. A crucial issue for the current models is that they usually underestimate the scavenging for convective events (probably due to phoretic effects and electrical effects, see above). Another reason is that the rain data are usually averaged in space and time.

The proposed correction consists in considering an aerosol of radius in  $[0.1 - 5] \mu\text{m}$  as an aerosol of radius  $5 \mu\text{m}$  once  $p_0$  is above a cut-off value (for instance  $25 \text{ mm h}^{-1}$ ). This brute-force approach is justified by the uncertainties in the microphysical data.

For instance with the parameterization of Baklanov and Sorensen (2001) (Fig. 8), a particle of radius 1 μm has a scavenging coefficient multiplied by a factor 3.5 above the threshold ( $\frac{3.71}{1.06}$ ). Notice that this is relevant with the difference between the coefficients  $A$  for the convective washout and rainout (Table 11,  $3.36 \times 10^{-4} / 8.4 \times 10^{-5} \simeq 4$ ).

3.3.4.4. Sub-grid parameterization of the rain rate.

The values of the precipitation intensity are mean values over a grid box. As for violent rains, this implies that the “true values” can be underestimated. Some parameterizations (which is not specific for radionuclides) use a local rain rate applied to the rainy fraction of the grid box. For instance, the simplest form is to replace the parameterization  $A p_0^B$  by  $f \times A(p_0/f)^B$  with  $f$  the rainy fraction (Rotstajn and Lohmann, 2002).  $f$  may be estimated on the basis of the cloud cover ( $cc$ ).

In Stohl et al. (2005), only  $p_0$  is modified as  $p_0/f$  with  $f = \max(0.05, cc \times F(p_0))$  and  $F(p_0)$  a function of the dynamic and convective precipitations.

3.3.4.5. Parameterization of snowfall. The scavenging by snow is poorly known and the measured data are usually not related to radionuclides. The uncertainties are related to the large variety of types and shapes of solid hydrometeors. In many works, it is usually assumed that the scavenging coefficient is the same one as for rain. The experimental data are indeed highly scattered. For instance, in Jylhä (2000),  $A$  is estimated below  $10^{-6}$  (for an application to the removal of sulphur emission in the vicinity of a power station), a weak value. In Sparmacher et al. (1993), the scavenging coefficients for snow are about five times larger than the washout coefficients for rain (Table 12). We refer to Sparmacher et al. (1993) for the discussion about the measurement uncertainties (due to thermophoresis and diffusiohoresis).

A commonly used parameterization is given in Table 13 (for instance Sofiev et al., 2006; Baklanov and Sorensen, 2001, following Maryon et al., 1992). The values have to be compared to those given in Table 11 for rain scavenging.

We can also refer to Rotstajn and Lohmann (2002) with a parameterization under the form

$$A = \frac{E\beta(T)F}{2\rho_{\text{snow}}}, \tag{28}$$

Table 12

Enhancement of the scavenging coefficient parameterized by  $A = A \times p_0^B$  for snowfall (fits from measurements, Sparmacher et al., 1993)

Particle diameter (in μm)	Washout (rain)	Snow
$d_p = 2.3 \times 10^{-1}$	$2.34 \times 10^{-7} p_0^{0.59}$	—
$d_p = 4.6 \times 10^{-1}$	$3.14 \times 10^{-7} p_0^{0.60}$	$1.6 \times 10^{-6} p_0^{0.62}$
$d_p = 9.8 \times 10^{-1}$	$2.56 \times 10^{-7} p_0^{0.94}$	$8.1 \times 10^{-7} p_0^{0.89}$
$d_p = 1.66$	—	$3.49 \times 10^{-6} p_0^{1.09}$
$d_p = 2.16$	$1.72 \times 10^{-6} p_0^{0.61}$	—

Table 13

Parameterizations of the NAME model for the scavenging by snow (Maryon et al., 1992)

Scavenging	$A$	$B$
Below-cloud	$8.05 \times 10^{-5}$	0.305
In-cloud (convective)	$3.35 \times 10^{-4}$	0.79
In-cloud (dynamic)	$8.05 \times 10^{-5}$	0.305

with  $E$  a collision efficiency for snow,  $\rho_{\text{snow}}$  the density for snow (typically  $100 \text{ kg m}^{-3}$ ) and  $F$  a flux of precipitation (in  $\text{kg m}^{-2} \text{ s}^{-1}$ ).  $\beta$  is the slope coefficient of a Marshall–Palmer distribution for snow, given as a decreasing function of temperature (Rotstajn, 1997). The collision efficiency is supposed to be much lower than for rain, which lowers  $A$  (while the large density of snow leads to an increase).

3.3.4.6. Washout ratio. Another possible description of the washout is provided by the so-called washout ratio

$$W = \frac{C_{\text{aq}}}{C_{\text{g}}}, \tag{29}$$

with  $C_{\text{aq}}$  the dissolved concentration (with respect to water mass) and  $C_{\text{g}}$  the gaseous concentration (with respect to air mass). Chamberlain (1991) gives the following values for iodine (p. 131) and cesium (p. 92):  $W_I \simeq 30$ ,  $W_{Cs} \simeq 500$ .

For gaseous iodine (cesium is bound on particles), one can try to link  $A$  and  $W$ . For the scavenging of a gas, the dissolved concentration at ground can be computed as in Seinfeld and Pandis (1998, p. 1005), p. 1005, Eq. (20)

$$(C_{\text{aq}})_a \simeq \frac{6K_c h}{U_{\text{drop}} D_{\text{f}}} C_{\text{g}}, \tag{30}$$

where  $K_c$  is the mass transfer coefficient and  $h$  the cloud height.  $U_{\text{drop}}$  (raindrop falling velocity) and  $D_r$  (representative diameter for the raindrops) have already been defined. Moreover,  $(C_{\text{aq}})_a = (\rho_{\text{water}}/\rho_{\text{air}})C_{\text{aq}}$ .

One has moreover

$$A = \frac{6 \times 10^{-3} p_0 K_c}{3600 U_{\text{drop}} D_r}, \quad (31)$$

with  $p_0$  in  $\text{mm h}^{-1}$ . As  $\rho_{\text{water}}/\rho_{\text{air}} \simeq 10^3$ , one gets

$$A \simeq \frac{W p_0}{3600 h}. \quad (32)$$

This formula is also given in a similar form in Underwood (2001, p. 43). For  $h \simeq 1000 \text{ m}$  and  $W_I \simeq 30$ , one gets  $A \simeq 10^{-5} p_0$ . The magnitude is coherent with the previous parameterizations.

Notice that the washout ratio of organic iodine is low, which justifies the weak scavenging (Table 10).

#### 4. Conclusion

We have reviewed the detailed and parameterized models of dry deposition and wet scavenging that can be used for radionuclides. A model hierarchy is available for the modellers, ranging from constant values (for instance for dry deposition velocities) to detailed microphysical modelling. Detailed models for particles are strongly related to the size distribution, which is usually poorly known even for the release.

For below-cloud scavenging, the parameterization of the scavenging coefficient  $A = A p_0^B$  can be justified. The main uncertainties are actually related to the rain intensity. A key point is that there is a strong coherence among all available parameterizations.

There are at least two complementary ways for improving the modelling of these processes.

At the microphysical level, many effects have still to be included. For instance, electrostatic and phoretic effects may strongly impact the submicronic distribution and may explain the underestimation of the scavenging coefficient. Specific orographic effects related to atmospheric dispersion (orographic enhancement of the wet deposition through the “seeder-feeder” effect, Fournier et al., 2005) have also to be included in comprehensive models. The description of a polydisperse distribution of aerosols may also strongly modify the current parameterizations, usually based on a monodisperse distribution. The hygroscopic characteristics of particles may also

impact these properties and are related to the chemical composition of particles to which radionuclides are bound (Chate et al., 2003). Notice that a key point is to have an aerosol model in which the radionuclides can be included.

A second approach is related to the meteorological fields and the way they are used. Using averaged quantities for rain intensities may underestimate the scavenging due to the underlying nonlinearities. Convective rains play a key role and are usually not well taken into account. Using data assimilation of measured data related to rain in dispersion models may be a challenging but promising task for improving the modelling.

Due to the large amount of uncertainties, one conclusion is that using ensemble forecast on the basis of a set of different parameterizations and of different meteorological data is required for modelling the dispersion of radionuclides (Galmarini et al., 2004; Mallet and Sportisse, 2006).

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